# PHOTOLYSIS OF SO<sub>2</sub> IN THE PRESENCE OF FOREIGN GASES III. QUENCHING OF EMISSION BY FOREIGN GASES

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SUMMARY

The emission yields of  $SO_2$  at 2.0 Torr pressure were measured in the presence of foreign gases at 25°C with incident radiation at 3130 Å. Experiments were performed without and with 0.33 Torr of NO present in order to measure the total and singlet relative emission efficiencies. The foreign gases studied were NO,  $CO_2$ , CO,  $N_2$ ,  $O_2$ ,  $N_2O$ ,  $H_2O$ , OCS, and thiophene. The main purpose of the study was to extend the quenching gas pressure range beyond that used previously, to see if deviations occurred from the Stern–Volmer quenching plots. Such deviations did occur for the emitting singlet state,  ${}^1SO_2$ , but not for the emitting triplet state,  ${}^3SO_2$ . It is concluded that  ${}^1SO_2$  is neither the state initially produced when the radiation is absorbed, nor the exclusive precursor state to  ${}^3SO_2$  formation. Relative quenching constants obtained for the singlet and triplet emitting states agree well with previous work. There is no excess  ${}^3SO_2$  produced over that predicted from Stern–Volmer quenching.

## INTRODUCTION

 $SO_2$  can be photoexcited at wavelengths between about 2500 and 3400 Å to give two emitting states, a singlet,  ${}^{1}SO_2$ , and a triplet,  ${}^{3}SO_2$ . The photophysics of the primary process has been well established to be ${}^{1-3}$ :

$^{1}SO_{2}$	$\rightarrow$ SO <sub>2</sub> + $h\nu_{\rm f}$	(1a)
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 $\rightarrow SO_2$  (1b)

 ${}^{1}\mathrm{SO}_{2} + \mathrm{SO}_{2} \rightarrow {}^{3}\mathrm{SO}_{2} + \mathrm{SO}_{2} \tag{2a}$ 

 $\rightarrow 2SO_2$  (2b)

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$$^{3}SO_{2} \rightarrow SO_{2} + h\nu_{p}$$
 (3a)

$$\rightarrow$$
 SO<sub>2</sub> (3b)

$${}^{3}\mathrm{SO}_{2} + \mathrm{SO}_{2} \to 2\mathrm{SO}_{2} \tag{4}$$

Sidebottom *et al.*<sup>4</sup> recently found the  ${}^{1}SO_{2}$  emission to have a non-exponential decay when excitation was at wavelengths below 2860 Å. Their data were consistent with the hypothesis of one emitting electronic state in which rate constants for loss depend on the vibrational energy. For incident radiation above 2980 Å, the singlet decay was exponential and could be attributed exclusively to low-lying vibrational levels of  ${}^{1}SO_{2}$ .

Evidence has accumulated, however, that another singlet level of SO<sub>2</sub> (or an isomer) is involved in the photochemistry. Cehelnik *et al.*<sup>5</sup> required such a state, which they called SO<sub>2</sub>\*, to explain CO<sub>2</sub> production in the presence of CO. They proposed that most collisional deactivation of  ${}^{1}SO_{2}$  produced this state. These ideas have been corroborated recently in Calvert's laboratory<sup>6,7</sup>.

 $^{1}SO_{2}$  is known to be quenched by gases other than  $SO_{2}$ <sup>8</sup>:

$$^{1}SO_{2} + NO \rightarrow removal$$
 (5)

$$^{1}SO_{2} + M \rightarrow removal$$
 (6)

We have considered NO separately because of its special utility to us, as discussed later. If the outlined mechanism is correct, then  ${}^{1}SO_{2}$  should follow Stern–Volmer quenching:

$${}^{1}Q^{-1} = (k_{1} + k_{2}[SO_{2}] + k_{5}[NO] + k_{6}[M])/k_{1a}$$
 (a)

where  ${}^{1}Q$  is the fluorescence emission yield from  ${}^{1}SO_{2}$ . Thus at constant SO<sub>2</sub> and NO pressures,  ${}^{1}Q^{-1}$  should vary linearly with the gas pressure of added gas, M.

On the other hand, the possibility exists that  $SO_2^*$  is really the state initially formed when radiation is absorbed, and that  ${}^1SO_2$  is produced from it. If so, and if  $SO_2^*$  can also be quenched, then the Stern-Volmer plots would deviate from linearity. The possibility that  $SO_2^*$  is really the initially formed state is intriguing, since the lifetime of  ${}^1SO_2$  at zero pressure is about 70 times longer than that computed from the integrated absorption coefficient<sup>2</sup>. Previous work has shown no deviation in the Stern-Volmer plots, but all the work was done at relatively low pressures (< 2 Torr). In this paper we extend the pressure range to see if deviations occur.

In order to explain the CO<sub>2</sub> production in the photolysis of SO<sub>2</sub>-CO mixtures, it is necessary that more triplet SO<sub>2</sub> be present than expected from the known quenching constants, *i.e.* there must be an additional source of triplet  $SO_2^{5-7, 9}$ . Such an effect was also observed in the SO<sub>2</sub>-sensitized phosphorescence of biacetyl, where foreign gases did not completely quench the biacetyl phosphorescence<sup>10, 11</sup>. In our laboratory we have interpreted these results with additional non-emitting triplet states of  $SO_2^{5, 9, 10}$  whereas the Calvert group has

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postulated an additional source for the emitting  ${}^{3}SO_{2}$  at high gas pressures<sup>6</sup>, <sup>7</sup>, <sup>11</sup>. In this paper, therefore, we have extended the quenching measurements using up to one atmosphere pressure of foreign gases to see if there remains a residual emission from  ${}^{3}SO_{2}$ .

# EXPERIMENTAL

Most of the emission system used has been described in a previous paper<sup>12</sup>. The major change was in the detection of the emission signal, which here was mechanically chopped with a 16 slot blade powered by a Princeton Applied Research (PAR) chopper, Model 125, to provide a chopping frequency of 333 Hz. The modulated emission signal passed through a Corning C.S. 0–52 filter to remove wavelengths shorter than 3400 Å before it was detected with an EMI 9558B photomultiplier. The signal voltage was developed across a 10 M $\Omega$  resistor, and amplified with a PAR, Model 112, preamplifier ( $\times$  100).

The signal was then fed into a PAR (Model 121) Lock-in Amplifier–Phase Detector which was tuned to the chopping frequency. The signal, in phase with the reference signal coming from the mechanical chopper, was read directly on the meter of the lock-in amplifier.

Fluorescence signals were corrected for variations of the intensity of the exiting beam, which was monitored by means of an RCA 935 Photodiode, and for the background signal due to scattered light.

The wavelength used for the irradiation was 3130 Å. For this purpose a 200 W high-pressure Hg arc (Illumination Industries Inc.) was used in conjunction with two 3130 Å interference filters (Baird Atomic Inc.).

Pressure measurements were made with a CD 25 Pace Pressure Transducer Indicator for the 0.01-5 Torr range. A dibutyl phthalate manometer was used for the 2-50 Torr range, and a Wallace and Tiernan absolute pressure gauge for the 50-800 Torr range.

All the following gases used were from the Matheson Company.  $N_2$  (prep. grade), and  $O_2$  (extra dry) were used without further purification. CO (chemically pure grade) was used after passing it through a glass wool trap immersed in liquid  $N_2$ . NO (technical grade) was distilled from  $-186^{\circ}$ C into  $-196^{\circ}$ C.  $N_2$ O and CO<sub>2</sub> (bone dry) were used after degassing at liquid  $N_2$  temperature. SO<sub>2</sub> (anhydrous) was distilled from  $-98^{\circ}$ C to  $-130^{\circ}$ C.

The OCS was bubbled through a concentrated NaOH solution and then passed through two drying tubes tightly packed with Drierite. The effluent was collected in a U trap immersed in liquid N<sub>2</sub>. The OCS was then allowed to distil slowly through a tube tightly packed with Ascarite into a storage bulb. The OCS was then degassed repeatedly at  $-196^{\circ}$ C.

The H<sub>2</sub>O used was taken from the tap and used after degassing. Thiophene (Sharpless Chem. Inc.) was purified by gas chromatography with a 10% tricresyl

phosphate on Chromosorb G (N.A.W.) column heated at  $100^{\circ}$ C, operated with a helium flow rate of 60 cm<sup>3</sup>/min.

## RESULTS

The emission of SO<sub>2</sub> was monitored with incident radiation at 3130 Å so that only low-lying vibrational levels of  ${}^{1}SO_{2}$  would be involved. Initially the relative emission yield, Q, was examined for  $2.00 \pm 0.03$  Torr of SO<sub>2</sub> in the presence of NO. NO was chosen since it is a known triplet quencher, and small amounts should remove the phosphorescence<sup>13</sup>.

The results are shown in Fig. 1. In the absence of NO,  $Q = Q_0 = 59$ . Only about 0.1 Torr of NO is required to eliminate the phosphorescence, yielding a relative singlet yield,  ${}^{1}Q_{0}$ , of 23.6. The emission is not further affected until [NO] exceeds 0.4 Torr, at which point further increases reduce the signal due to quenching of  ${}^{1}SO_{2}$ . Thus with 0.33 Torr of NO,  ${}^{3}SO_{2}$  can be removed without significantly reducing  ${}^{1}SO_{2}$  emission. All additional runs to study the singlet emission alone were carried out with 0.33  $\pm$  0.03 Torr of NO and 2.0  $\pm$  0.03 Torr of SO<sub>2</sub> present.

The Stern–Volmer plot for NO is shown in Fig. 2. Initially, at very low NO pressures, the reciprocal relative emission yield,  $Q^{-1}$ , rises with the NO pressure as the triplet emission is quenched, then follows a linear law. However, for pressures



Fig. 1. Log-log plot of the relative emission yield at  $25^{\circ}$  C vs. 0.1 + [NO] for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr.



Fig. 2. Plot of reciprocal relative emission yield at 25° C vs. the NO pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr.

of NO above about 6 Torr, the emission falls more rapidly than expected from an extrapolation of the low pressure results.

Figures 3-10 show Stern-Volmer plots for 8 different quenching gases, M. In each Figure there is a plot for data obtained in the absence of NO as well as in the presence of 0.33 Torr of NO. The data are plotted as the ratio of the emission yield in the absence of M ( $Q_0$  with NO absent,  ${}^1Q_0$  with NO present) to that in the presence of M (Q with NO absent,  ${}^1Q$  with NO present). In all cases replicate runs were made.

In the presence of NO, a direct measure is being made of the quenching of the singlet emission from  ${}^{1}SO_{2}$ . In every case the plot curves upward as the pressure is enhanced. This is particularly marked with CO and N<sub>2</sub> where measurements could be made to relatively high pressures, because these gases are relatively inefficient quenchers. For these gases, the emission at 30 Torr pressure is one-half that expected from an extrapolation of the low pressure asymptotes.

In the absence of NO, the total emission yield from both  ${}^{1}SO_{2}$  and  ${}^{3}SO_{2}$  is being measured. There is no reason why a Stern–Volmer plot should be obeyed for



Fig. 3. Plots of the ratio of the emission yield at  $25^{\circ}$ C in the absence of CO<sub>2</sub> to that in its presence vs. the CO<sub>2</sub> pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and  ${}^{1}Q_{0}/{}^{1}Q$ ; the computed curve for  ${}^{3}Q_{0}/{}^{3}Q$  is also shown. The ordinates for  ${}^{1}Q_{0}/{}^{1}Q$  and  ${}^{3}Q_{0}/{}^{3}Q$  are displaced for clarity. In all cases the intercept is unity.

a composite emission, but it is convenient to present the data in this way. In fact, if the quenching efficiencies compared to  $SO_2$  are comparable for the two states, the plot will not show the break expected for two-state quenching. Only in the case of thiophene (Fig. 10) is the relative efficiency for the two states markedly different. Thiophene is a very good quencher for  ${}^{3}SO_2$ , as shown by the fact that only 0.060 Torr of thiophene is needed to completely quench the triplet emission, and the break in the curve is obvious. Again all the curves, except that for  $H_2O$ , show the upward curvature as the pressure is increased, although the deviation from linearity is not as marked as when NO is present. For each of the gases except  $H_2O$  and thiophene, measurements were made to one atmosphere pressure and in every case the emission was reduced below the limit of detection ( $Q_0/Q > 100$ ).

Two of the quenching gases gave rise to unique complications.  $O_2$  reacts with NO to produce NO<sub>2</sub>. Consequently pressures of  $O_2$  greater than 20 Torr could not be used with NO present. For H<sub>2</sub>O at pressures above 12 Torr, the emission signal was not reduced with the addition of more H<sub>2</sub>O vapor. This



Fig. 4. Plots of the ratio of the emission yield at  $25^{\circ}$  C in the absence of CO to that in its presence  $\nu s$ , the CO pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and  ${}^{1}Q_{0}/{}^{1}Q$ ; the computed curve for  ${}^{3}Q_{0}/{}^{3}Q$  is also shown. The ordinates for  ${}^{1}Q_{0}/{}^{1}Q$  and  ${}^{3}Q_{0}/{}^{3}Q$  are displaced for clarity. In all cases the intercept is unity.

residual emission was attributed to light scattering by aerosol formation. This conclusion was supported by the finding that adding an atmosphere of air did not affect the residual signal. In the absence of  $H_2O$  vapor, an atmosphere of air completely suppresses the signal.

For thiophene the emission yields in the presence or absence of NO are virtually identical showing that thiophene efficiently removes the triplet emission. For the other seven quenching gases, the triplet yield can be computed from the smooth curves fitting the data for  $Q_0/Q$  and  ${}^1Q_0/{}^1Q$ . From these curves and the values of  $Q_0 = 59$  and  ${}^1Q_0 = 23.6$ , the values of  ${}^3Q_0/{}^3Q \equiv (Q_0 - {}^1Q_0)/(Q - {}^1Q)$  can be computed. These computed curves are shown in Figs. 3-9. They are linear in all cases except possibly for OCS which may show a slight downward trend. Thus the triplet yields follow a simple Stern-Volmer quenching law, whereas the singlet yields do not.



Fig. 5. Plots of the ratio of the emission yield at  $25^{\circ}$ C in the absence of N<sub>2</sub> to that in its presence vs. the N<sub>2</sub> pressure for incident radiation at 3130 Å and [SO<sub>2</sub>] = 2.0 Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and  ${}^{1}Q_{0}/{}^{1}Q$ ; the computed curve for  ${}^{3}Q_{0}/{}^{3}Q$  is also shown. The ordinates for  ${}^{1}Q_{0}/{}^{1}Q$  and  ${}^{3}Q_{0}/{}^{3}Q$  are displaced for clarity. In all cases the intercept is unity.

### DISCUSSION

# <sup>1</sup>SO<sub>2</sub> emission

The deviation from linearity in the Stern–Volmer plots (Figs. 3–10) indicates that the emitting state is not the state which absorbs radiation. Since it is known<sup>4</sup> that, with incident radiation at 3130 Å, the emission comes from low-lying vibrational levels of <sup>1</sup>SO<sub>2</sub>, there are two possible explanations which need to be considered. Either absorption occurs to high vibrational levels of <sup>1</sup>SO<sub>2</sub> and the vibrational energy is removed by collision prior to emission, or the absorption is to a different



Fig. 6. Plots of the ratio of the emission yield at 25°C in the absence of O<sub>2</sub> to that in its presence vs. the O<sub>2</sub> pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and  ${}^{1}Q_{0}/{}^{1}Q$ ; the computed curve for  ${}^{3}Q_{0}/{}^{3}Q$  is also shown. The ordinates for  ${}^{1}Q_{0}/{}^{1}Q$  and  ${}^{3}Q_{0}/{}^{3}Q$  are displaced for clarity. In all cases the intercept is unity.

electronic level. In the former case the Stern-Volmer plot should not display two non-zero quenching regions as the pressure is raised. Consequently, the latter explanation must be correct. The simplest mechanism to explain the data is:

$$SO_2 + h\nu \rightarrow SO_2^*$$
 (7)

$$SO_2^* \rightarrow {}^1SO_2$$
 (8), (-8)

$$\mathrm{SO}_2^* + \mathrm{SO}_2 \to \mathrm{removal}$$
 (9)

$$SO_2^* + NO \rightarrow removal$$
 (10)

$$SO_2^* + M \rightarrow removal$$
 (11)



Fig. 7. Plots of the ratio of the emission yield at  $25^{\circ}$  C in the absence of N<sub>2</sub>O to that in its presence vs. the N<sub>2</sub>O pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and <sup>1</sup>Q<sub>0</sub>/<sup>1</sup>Q; the computed curve for <sup>3</sup>Q<sub>0</sub>/<sup>3</sup>Q is also shown. The ordinates for <sup>1</sup>Q<sub>0</sub>/<sup>1</sup>Q and <sup>3</sup>Q<sub>0</sub>/<sup>3</sup>Q are displaced for clarity. In all cases the intercept is unity.

$${}^{1}\mathrm{SO}_{2} \qquad \rightarrow \mathrm{SO}_{2} + h\nu_{\mathrm{f}} \tag{1a}$$

$$\rightarrow$$
 SO<sub>2</sub> (1b)

$${}^{1}\mathrm{SO}_{2} + \mathrm{SO}_{2} \rightarrow \mathrm{removal} \tag{2}$$

$${}^{1}\mathrm{SO}_{2} + \mathrm{NO} \rightarrow \mathrm{removal}$$
 (5)

$$^{1}\text{SO}_{2} + \text{M} \rightarrow \text{removal}$$
 (6)

The state initially formed by absorption,  $SO_2^*$ , has an emission lifetime of about  $6 \times 10^{-7}$  s as computed from the integrated absorption coefficient<sup>2</sup>. Since emission is not seen from  $SO_2^*$ ,  $k_8$  must be  $\ge 10^8 \text{ s}^{-1}$ , which would then give a lifetime of  $SO_2^* < 10^{-8}$  s. This value is lower than the  $50 \times 10^{-9}$  s limiting lifetime that could have been seen by Sidebottom *et al.*<sup>4</sup>.



Fig. 8. Plots of the ratio of the emission yield at  $25^{\circ}$ C in the absence of H<sub>2</sub>O to that in its presence vs. the H<sub>2</sub>O pressure for incident radiation at 3130 Å and [SO<sub>2</sub>] = 2.0 Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and <sup>1</sup>Q<sub>0</sub>/<sup>1</sup>Q; the computed curve for <sup>3</sup>Q<sub>0</sub>/<sup>3</sup>Q is also shown. The ordinates for <sup>1</sup>Q<sub>0</sub>/<sup>1</sup>Q and <sup>3</sup>Q<sub>0</sub>/<sup>3</sup>Q are displaced for clarity. In all cases the intercept is unity.

The mechanism predicts the rate law:

$${}^{1}Q_{0}/{}^{1}Q = \frac{(k_{11}k_{-8} + (1+a)k_{6}k_{8} + k_{11}\beta + k_{6}k_{11}[M])[M]}{k_{8}\{(1+a)\beta + k_{-8}a\}} + 1 \qquad (b)$$

where  $\alpha \equiv (k_9[SO_2] + k_{10}[NO])/k_8$ 

$$\beta \equiv k_1 + k_2[\mathrm{SO}_2] + k_5[\mathrm{NO}]$$

and  ${}^{1}Q_{0}$  and  ${}^{1}Q$  are the singlet emission yields in the absence and presence of quenching gas respectively.

Under our conditions,  $\alpha \ll 1$ ,  $\beta \gg k_{-8}$ ,  $k_1 \ll k_2[SO_2]$ , and  $k_{11}/k_8 < k_6/\beta$ . Thus the limiting low pressure slopes of the Stern–Volmer plots for  ${}^1Q_0/{}^1Q$  are approximately equal to  $k_6/\beta$ . From the values of  $k_6/\beta$ ,  $k_6/k_2$  can be computed, and they are listed and compared with literature values in Table 1. The uncertainty in these values is  $\pm 10-20\%$ . Our values for NO and CO<sub>2</sub> are somewhat higher than found by Mettee<sup>8</sup>, but our value for CO<sub>2</sub> is in excellent agreement with that of



Fig. 9. Plots of the ratio of the emission yield at  $25^{\circ}$ C in the absence of OCS to that in its presence vs. the OCS pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The triplet yields are computed from the smooth curves passing through the data for Q<sub>0</sub>/Q and  ${}^{1}Q_{0}/{}^{1}Q$ ; the computed curve for  ${}^{3}Q_{0}/{}^{3}Q$  is also shown. The ordinates for  ${}^{1}Q_{0}/{}^{1}Q$  and  ${}^{3}Q_{0}/{}^{3}Q$  are displaced for clarity. In all cases the intercept is unity.

Rao *et al.*<sup>14</sup>. For CO and N<sub>2</sub> our results are in excellent agreement with the literature values. For O<sub>2</sub> our value is somewhat high and probably erroneous because of the reaction of NO and O<sub>2</sub> in our system.

# <sup>3</sup>SO<sub>2</sub> emission

The traditional view is that  ${}^{3}SO_{2}$  is produced from  ${}^{1}SO_{2}$ . Since under our conditions the first-order decay processes (reactions 1 and 3) are unimportant, the additional steps needed to explain the  ${}^{3}SO_{2}$  emission are:

$1SO_2 + SO_3$	$_2 \rightarrow \ ^3\mathrm{SO}_2 + \mathrm{SO}_2$	(2a)
	$\rightarrow 2SO_2$	(2b)
$^{1}SO_{2} + M$	$\rightarrow {}^{3}SO_{2} + M$	(6a)
	$\rightarrow$ SO <sub>2</sub> + M	(6b)
${}^{3}SO_{2} + M$	$\rightarrow$ SO <sub>2</sub> + M	(12)

## TABLE 1

## SUMMARY OF RATE CONSTANT DATA

М	k <sub>4</sub> /k <sub>2</sub>			k12/k4		
	This work	Metteeª	Calvert	This work	Mettee <sup>a</sup>	Calvert
NO	0.86	0.59		~ 100	64	190ª
$CO_2$	0.76	0.63	0.73 <sup>b</sup>	0.55	0.31	0.29ª
CO	0.32	0.35		0.26	0.30	0.22°
$N_2$	0.25	0.28	0.29°	0.23	0.27	0.25ª
O2	0.37	0.27	0.31ъ	0.42	0.37	0.25ª
N <sub>2</sub> O	0.56	_	_	0.38	_	_
H <sub>2</sub> O	0.93	—		1.8	_	2.28ª
OCS	0.72	_		0.60		_
Thiophene	2.4 2.0 <sup>r</sup>	-	—	≥100	—	_

<sup>a</sup> Ref. 8.

<sup>b</sup> Ref. 14.

° Ref. 15.

<sup>d</sup> Ref. 13; direct excitation of triplet at 3828 Å.

e Ref. 16; direct excitation of triplet 3828 Å.

<sup>f</sup> From data with NO absent.

At low pressures the collisional quenching of  $SO_2^*$  is unimportant, and the approximate rate law for  ${}^3SO_2$  becomes:

$${}^{3}Q_{0}/{}^{3}Q = \frac{(1 + k_{6}[M]/k_{2}[SO_{2}])(1 + k_{12}[M]/k_{4}[SO_{2}])}{(1 + k_{6a}[M]/k_{2a}[SO_{2}])}$$
(c)

Both Mettee<sup>8</sup> and we have found the Stern-Volmer plot to be satisfied at low quenching gas pressures, so that  $k_{2a}/k_2$  must be similar to  $k_{6a}/k_6$  for each quenching gas for this mechanism to apply. This is highly unlikely. Furthermore the fact that the Stern-Volmer plots for triplet emission do not show the curvature typical of singlet emission excludes the possibility that the triplet state is produced exclusively from the emitting singlet state.

A simple, but not necessarily correct, explanation of the results is that  ${}^{3}SO_{2}$  is produced directly by absorption:

$$SO_2 + hv \to {}^3SO_2 \tag{13}$$



Fig. 10. Plots of the ratio of the emission yield at  $25^{\circ}$ C in the absence of thiophene to that in its presence vs. the thiophene pressure for incident radiation at 3130 Å and  $[SO_2] = 2.0$  Torr. The circles are for data in the absence of NO and correspond to total emission yields. The squares are for data with 0.33 Torr of NO present and correspond to singlet emission yields. The different point notations indicate replicate runs. The ordinate for  ${}^{1}Q_{0}/{}^{1}Q$  is displaced for clarity; its intercept is unity.

with the contribution from reactions (2a) and (6a) being negligible. The rate law then becomes:

$${}^{3}Q_{0}/{}^{3}Q = 1 + k_{12}[M]/k_{4}[SO_{2}]$$
 (d)

A simple straight-line Stern-Volmer plot is predicted in conformance with our findings. The slopes of the plots give  $k_{12}/k_4$ [SO<sub>2</sub>], from which  $k_{12}/k_4$  can be computed. The computed values are listed in Table 1. Again the uncertainty is  $\pm 10-20\%$ . Except for CO<sub>2</sub>, and possibly O<sub>2</sub>, they agree reasonably well with those obtained by Mettee<sup>8</sup> in the same way, and by Calvert and his coworkers<sup>13, 16</sup> who directly excited <sup>3</sup>SO<sub>2</sub> with 3828 Å radiation. As mentioned before, our value

for  $O_2$  is probably unreliable because of the reaction between  $O_2$  and NO. It is not clear why our value for  $CO_2$  is high.

The data for the triplet yields show no excess production of triplet at high pressures. For example, with added nitrogen or CO the Calvert interpretation requires that  ${}^{3}Q_{0}/{}^{3}Q$  should increase by no more than a factor of two or possibly three. In fact, from our data an increase of at least a factor of 10 is computed.

The lack of excess  ${}^{3}SO_{2}$  production is more dramatically shown by the quenching of the total emission. In every case but with H<sub>2</sub>O, which produces an aerosol at pressures > 12 Torr, the total emission Q<sub>0</sub>/Q shows the upward curvature as the pressure is raised, but the curvature is less pronounced than that for  ${}^{1}Q_{0}/{}^{1}Q$ . Thus the triplet yield must show less curvature than the singlet yield. For all the quenching gases except H<sub>2</sub>O and thiophene, the results were extended to 1 atm pressure and the signal vanished, *i.e.* Q<sub>0</sub>/Q > 100. Since our value for Q<sub>0</sub>/ ${}^{1}Q = 2.5$ , the Calvert interpretation would require that Q<sub>0</sub>/Q not increase by more than a factor of 6 and probably less. It is clear that this interpretation is inconsistent with the observations.

#### ACKNOWLEDGEMENTS

We are deeply indebted to Professor Jack Calvert and his research group at Ohio State University for numerous stimulating discussions and for access to their data prior to publication. This work was supported through the Atmospheric Sciences Section of the National Science Foundation under Grant No. GA-33446X and the Environmental Protection Agency under Grant No. 800874, for which we are grateful. We also wish to thank Dr. Eckart Mathias for his help.

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